What can VLBI do for your research? The EVN and JIVE

By

Dr. Francisco Colomer

Joint Institute for VLBI ERIC (JIVE)

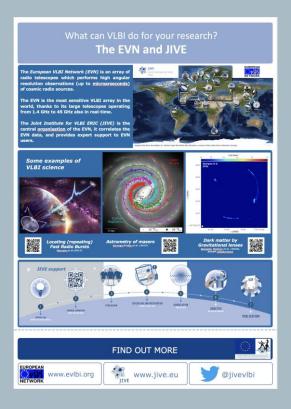


EAVN @ PyonengChang (Korea), 4-7 September 2018

Contents

- The European VLBI Network (EVN)
- The Joint Institute for VLBI ERIC (JIVE)
- (A few) recent VLBI results
- Globalization of VLBI

See also poster:





European VLBI Network (EVN)

- The EVN is a network of radio telescopes located primarily in Europe and Asia, with additional antennas in South Africa and Puerto Rico, which performs high angular resolution observations of cosmic radio sources.
- It is **the most sensitive VLBI array in the world**, thanks to the collection of extremely large telescopes that contribute to the network, operating from 1.4 GHz to 45 GHz (some up to 90 GHz), also in **real-time** (*e*EVN).
- The Joint Institute for VLBI ERIC (JIVE) correlates the EVN data and provides expert support to EVN users.



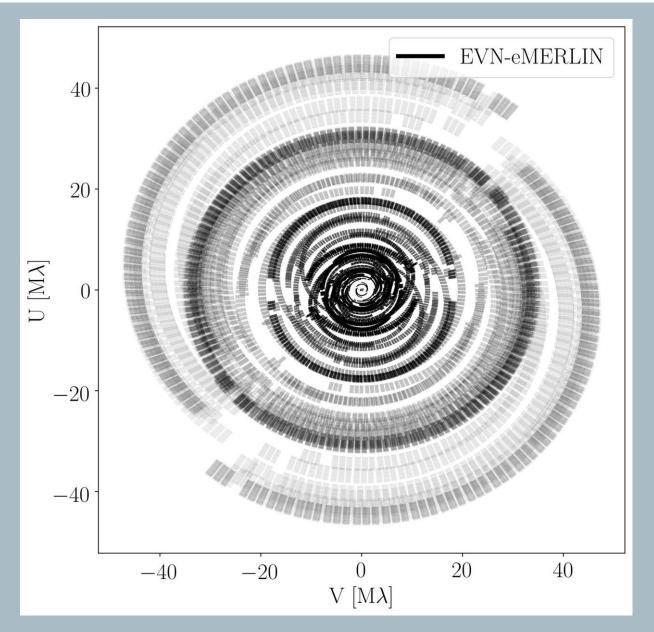
http://www.evlbi.org/





www.jive.eu / www.evlbi.org

EVN+MERLIN







Joint Institute for VLBI ERIC (JIVE)

An European Research Infrastructure Consortium (ERIC)

- 6 partner countries: NL (host), FR, ES, UK, SE, LV
- 4 associated institutions: INAF (IT), NRF (SA), MPIfR (DE), NAOC (Cn)







JIVE and EVN

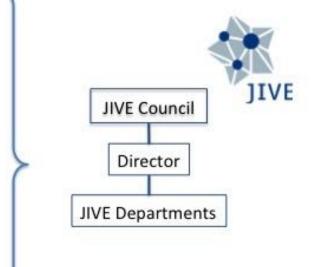
EUROPEAN VLBI NETWORK (EVN)

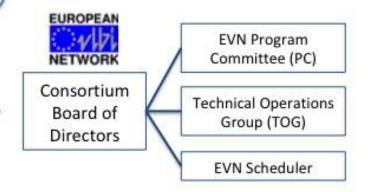
JOINT INSTITUTE FOR VLBI ERIC (JIVE)

The Netherlands (WSRT)
United Kingdom (JBO)
Spain (IGN-Yebes)
Sweden (OSO)
Latvia (VIRAC)
France (-)

Germany (E5)
Italy (Nt, Mc, SRT)
South Africa (HRAO, AVN)
China (SHAO, Urumqi)

Russia (KVAZAR) Finland (Metsahovi) Poland (Torun) USA (Puerto Rico, NASA/DSN) South Korea (KVN)









Joint Institute for VLBI ERIC (JIVE)

- Supports the European VLBI Network
 - Radio astronomy at its highest angular resolution
 - operations
 - correlation
 - data acquisition
- Research & development
 - Software (CASA 5.3)
 - hardware
 - e-VLBI
- Science support
- Training



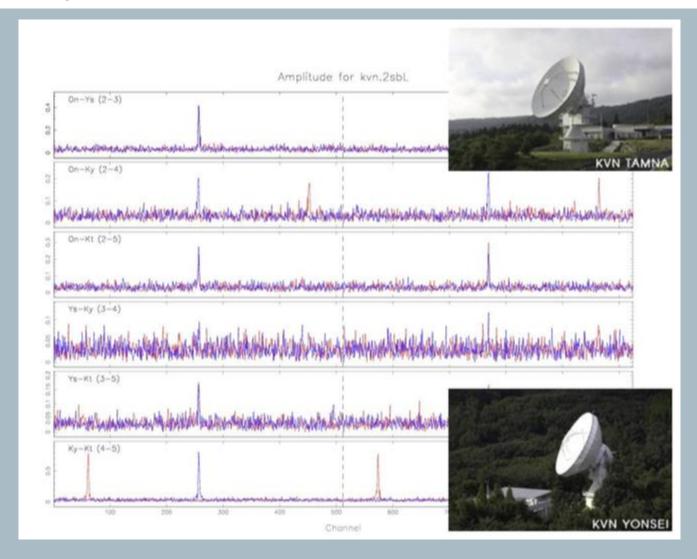
JIVE headquarters in Dwingeloo, the Netherlands.



http://www.jive.eu/



EVN/JIVE collaboration with KVN

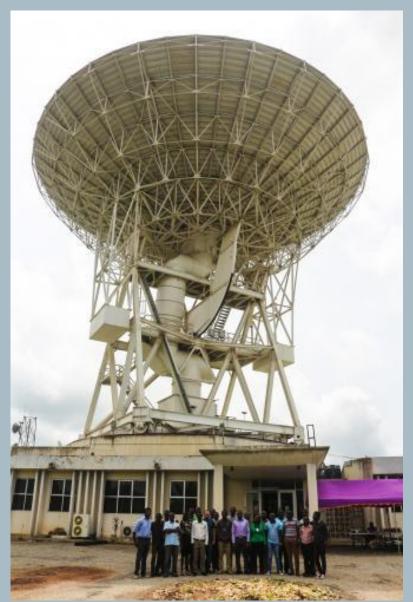


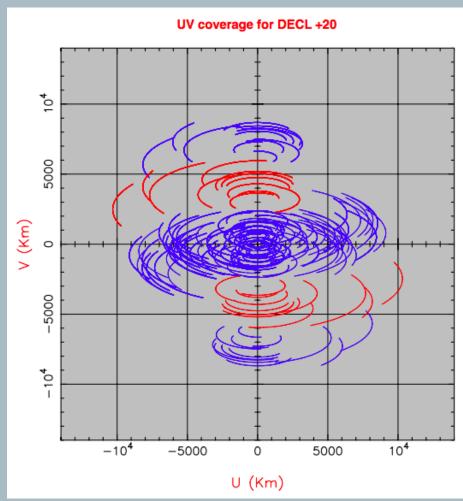
First e-VLBI fringes between EVN antennas (Onsala and Yebes) and two stations of the Korean VLBI Network at 22 GHz.





New VLBI station in Kutunse (Ghana)



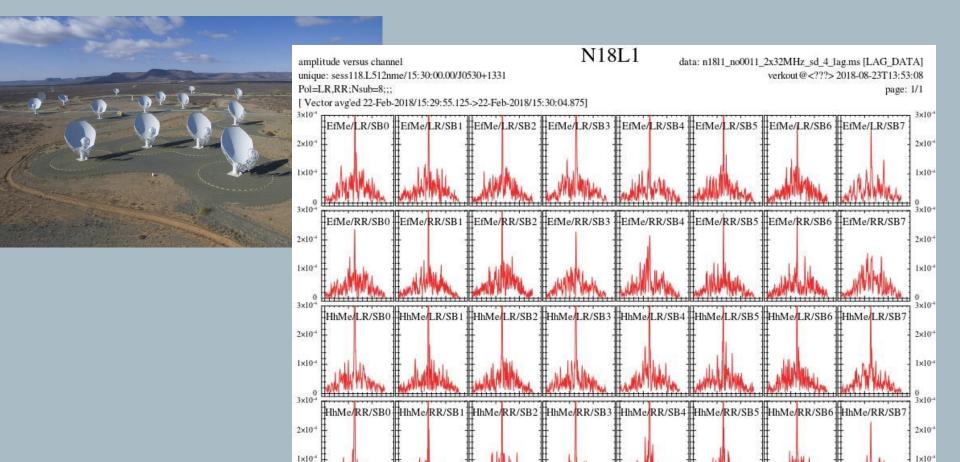


HART
GHANA
TORUN
YEBES40M
MEDICINA
WSTRBORK
JODRELL1
SHANGHAI
NOTO
GOONHILL
BADARY
ZELENCHK
SVETLOE
URUMQI





First VLBI fringes with MeerKAT!







New potential EVN partners



Haopin, 40 m, China



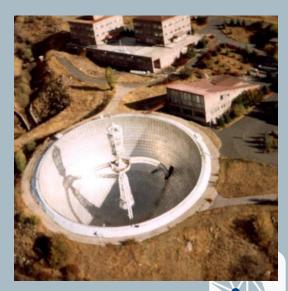
Zolochiv, 32 m, Ukraine



Azores, 32 m, Portugal



Usuda, 64 m, Japan

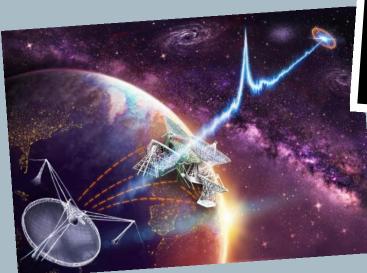


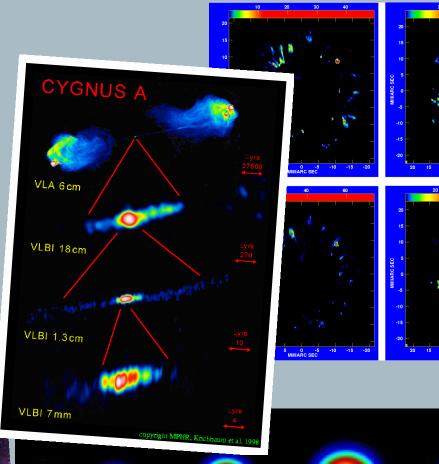
Aragats,50 m, Armenia



VLBI science highlights

- AGN, galaxies
- Star formation
- Evolved stars
- The transient Universe
- Astrometry
- Reference frames ...

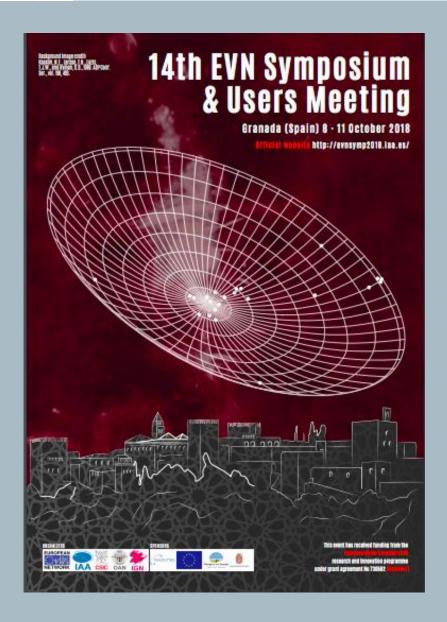








EVN Symposium 2018



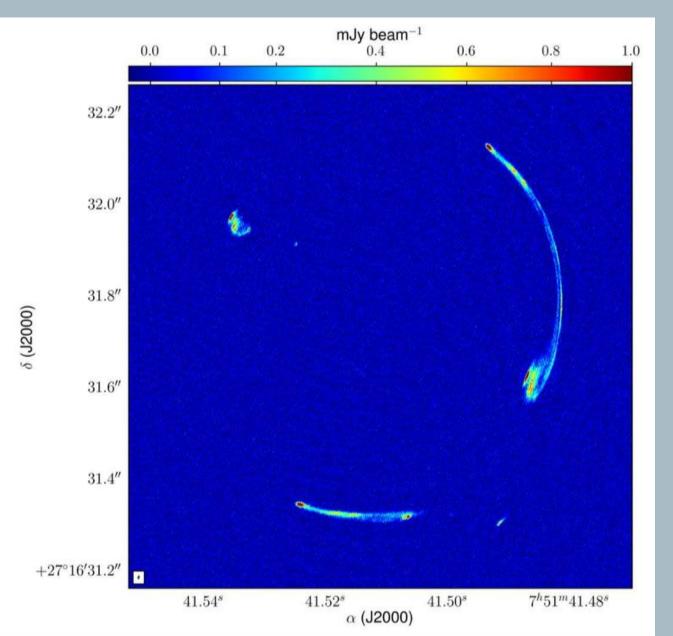
In Granada (Spain), October 8-11 2018

http://evnsymp2018.iaa.es/





Gravitational lenses and dark matter



"SHARP – V. Modelling gravitationally-lensed radio arcs imaged with global VLBI observations".

C. Spingola et al.

Monthly Notices of the Royal Astronomical Society, Volume 478, Issue 4, 21 August 2018, Pages 4816– 4829, https://doi.org/10.1093/mnr as/sty1326

JIVE PR:

http://www.jive.eu/new-imagessuper-telescope-bringastronomers-step-closerunderstanding-dark-matter

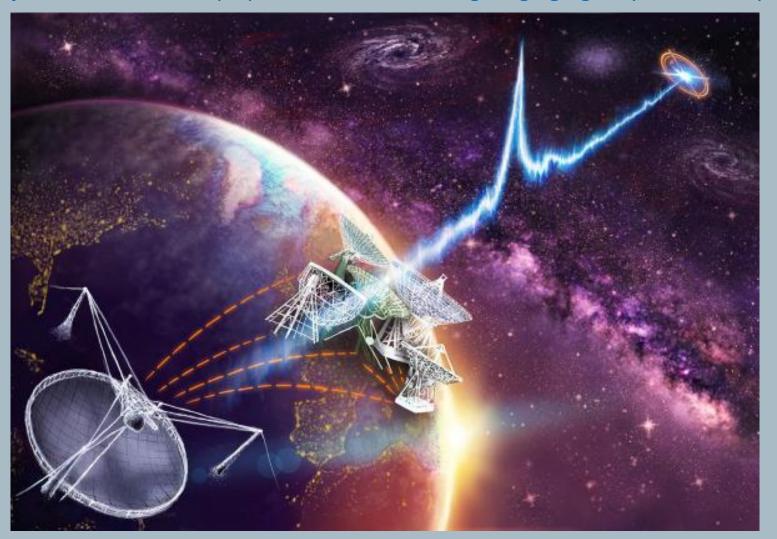




Locating mysterious Fast Radio Bursts

EVN enables the identification of the host galaxy.

http://jive.eu/astronomers-pinpoint-radio-flashes-long-long-ago-galaxy-far-far-away

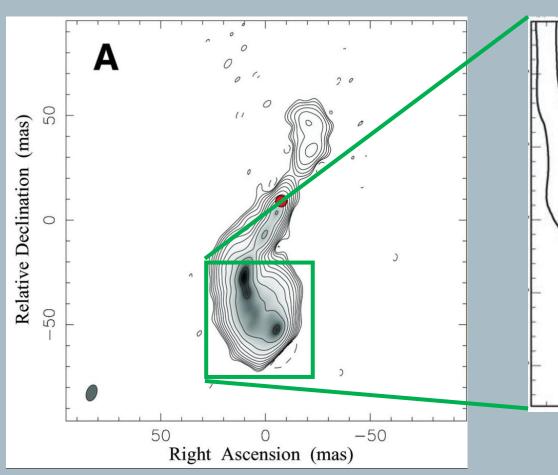


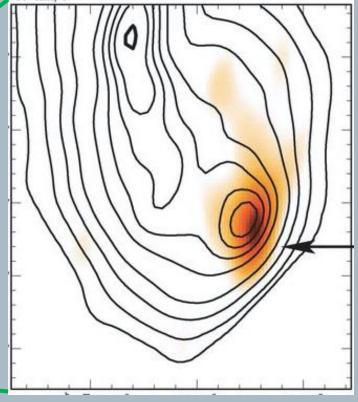




Material ejected from (or falling into) black holes.

http://science.sciencemag.org/content/341/6150/1082

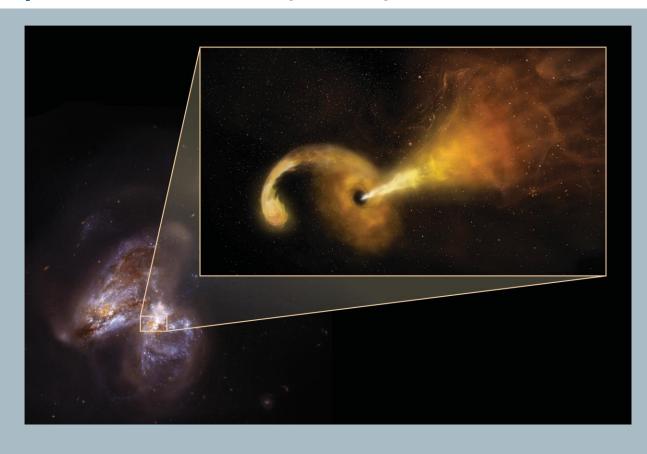






Tidal Disruption Event (TDE)

A supernova explosion was actually a star being pulled apart by a supermassive black hole: the powerful gravity of SMBH rips apart a star that has wandered too close.

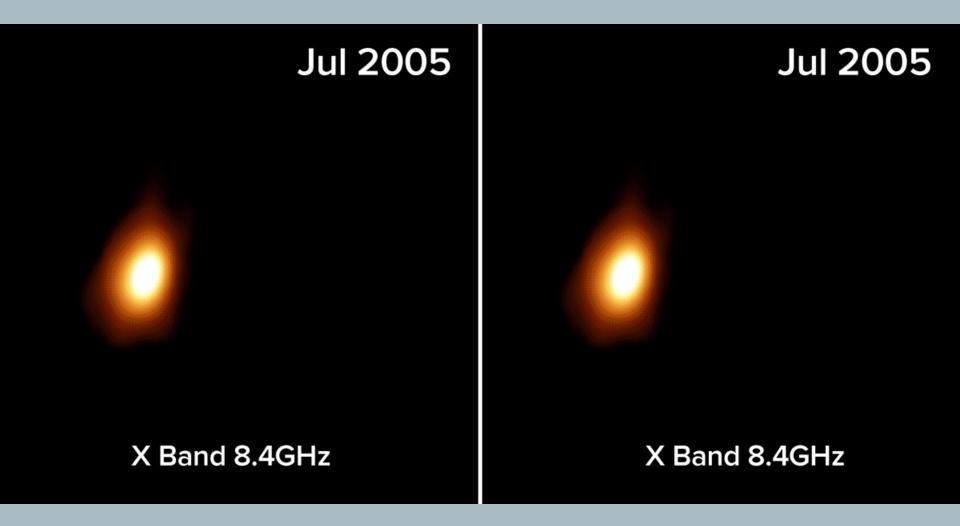


Mattila, S., Pérez-Torres, M., et al. 2018. A dust enshrouded tidal disruption event with a resolved radio jet in a galaxy merger. Science. DOI: 10.1126/science.aao4669





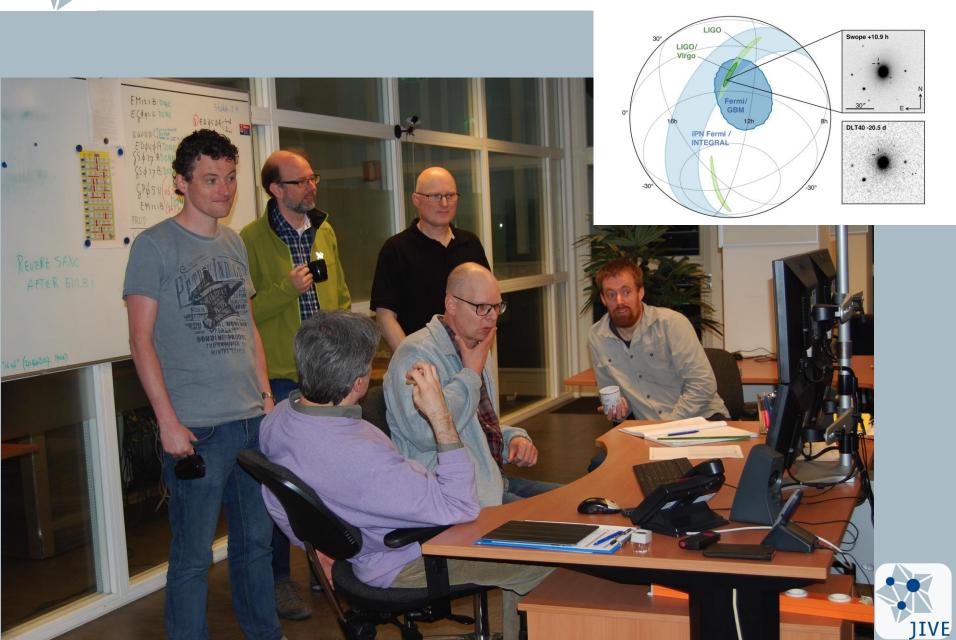
Tidal Disruption Event (TDE)







First EVN observation of a Gravitational Wave counterpart

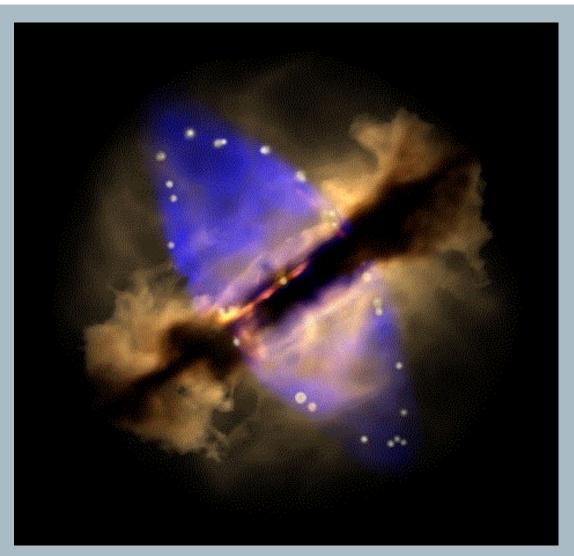




Birth of massive stars

EVN can trace gas motions around birth place.

http://jive.eu/w75n-outflow-onset

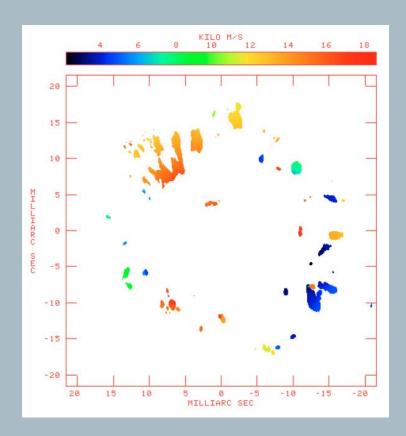






Envelopes of evolved stars

VLBI displays the star CSE evolution in real time.

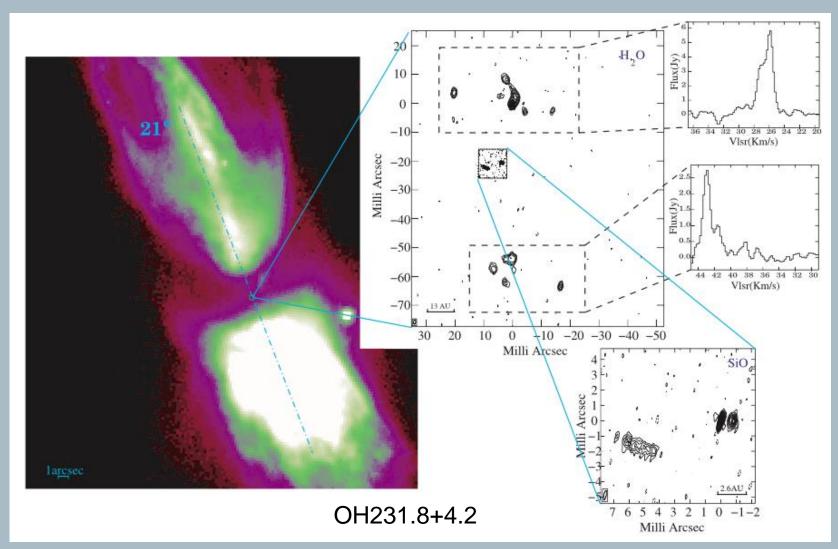








Mathematical Appears of the English How does axial asymmetry appears?



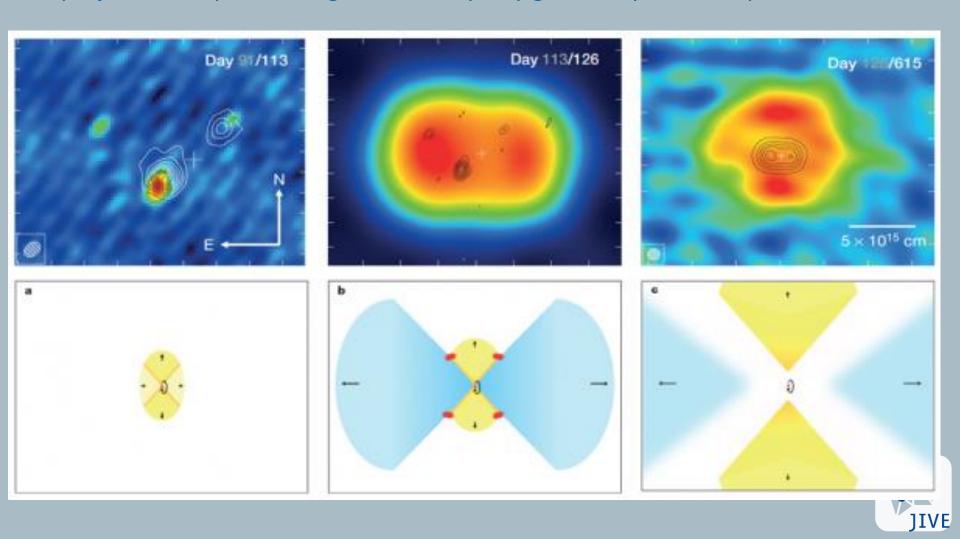




Nova Mon: stellar evolution

EVN solves riddle on gamma ray production.

http://jive.eu/sharp-radio-images-unravel-mystery-gamma-rays-stellar-explosions





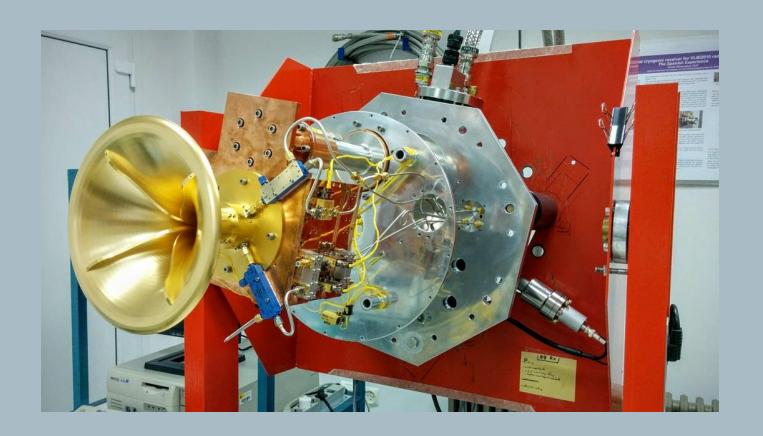
What can VLBI do for you?

FUTURE PROSPECTS



RADIONET BRAND receiver for EVN

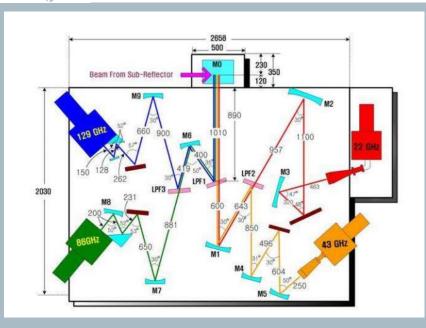
- Broadband receiver to cover 1.5-15.5 GHz.
- Allows simultaneous registration of masers of OH (1.6, 1.7, 4.7 and 6.0 GHz), and CH₃OH (6.7 and 12.2 GHz).







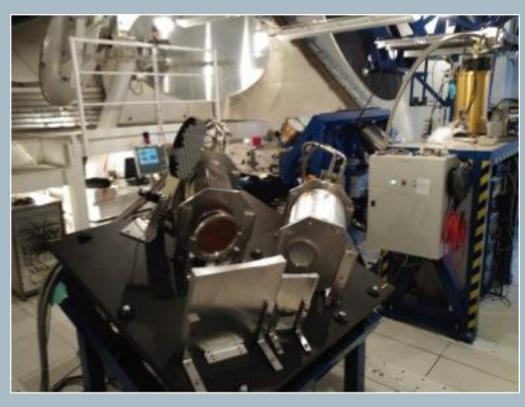
Simultaneous multi-frequency observations



K band (21-26 GHz) Q band (41-49 GHz) W band (76-100 GHz)

Water masers
SiO masers
HCN masers
Some methanol masers

Allows alignment of maser maps!





K/Q/W at IGN-Yebes 40-m RT





Compact K/Q/W bands receiver





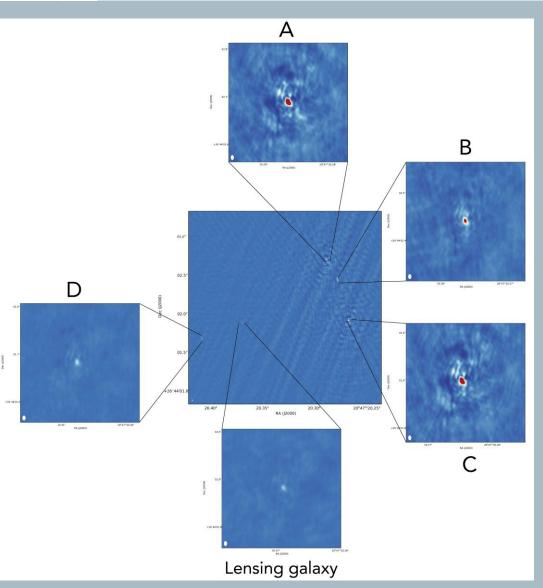
Also allows calibration of high frequency data (up to W band) using lower frequency (Q, K) simultaneous observations

See talks by Dr. Han, and by Maria Rioja!





Software highlights



Fringe Fitting for VLBI in CASA reaches maturity

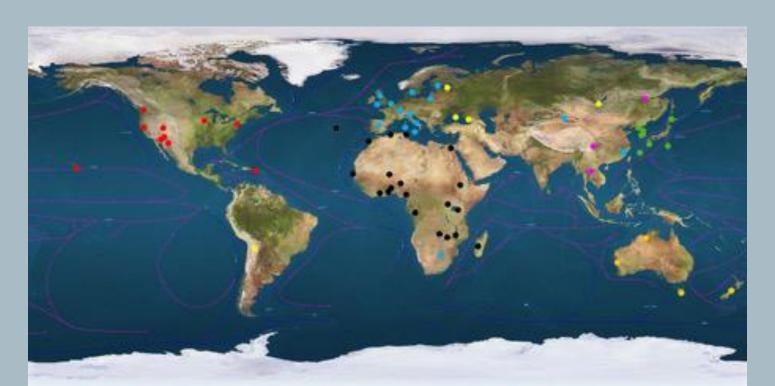
- Two very successful workshops to exercise new capability
- First image of a VLBI science target fully processed with the new CASA tools





Towards global VLBI

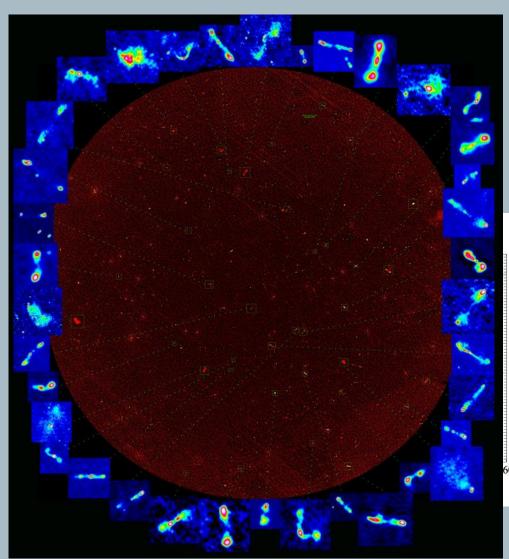
- "Earth VLBI alliance" concept being developed
 - EVN, VLBA, LBA, GMVA, EAVN, KVN, VERA, AVN...
 - Need to set a Global VLBI Working Group (GVWG) for coordination
- When SKA_1 is built, long baselines and northern hemisphere sky coverage will be needed!



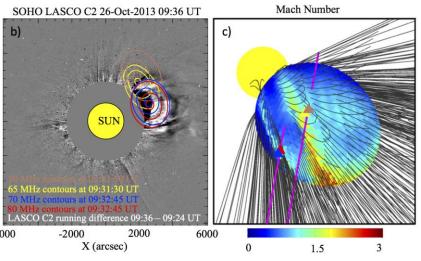




EVN/JIVE - LOFAR/ILT synergies











VLBI in the era of SKA

- The Square Kilometer Array (SKA) will be built in two phases.
- SKA-1 is 10% of the full SKA, and will lack very long baselines, which are provided by VLBI.









EC H2020 JUMPING JIVE



WP10: SKA-VLBI

- Explores the synergies between VLBI and SKA
- SKA1-MID shares
 frequencies with EVN;
 combined observations
 will provide
 unprecedented sensitivity.
- To develop a global VLBI
 Science Case including
 precision astrometry, large
 field-of-view VLBI, VLBI
 surveys and transients, etc.







- VLBI is the astronomical technique with highest angular resolution
- VLBI provides a huge add-on value to national RA facilities
- EVN is the most sensitive VLBI network in the world, open to all astronomers
- Members of EVN/JIVE are involved in cutting-edge research and development
- Development of new VLBI stations will improve image quality, and create local expert communities
- JUMPING JIVE fosters global VLBI need to set a Global VLBI Working Group (GVWG)
- e-VLBI networks are pathfinders to SKA





Thank you!

Francisco Colomer colomer@jive.eu

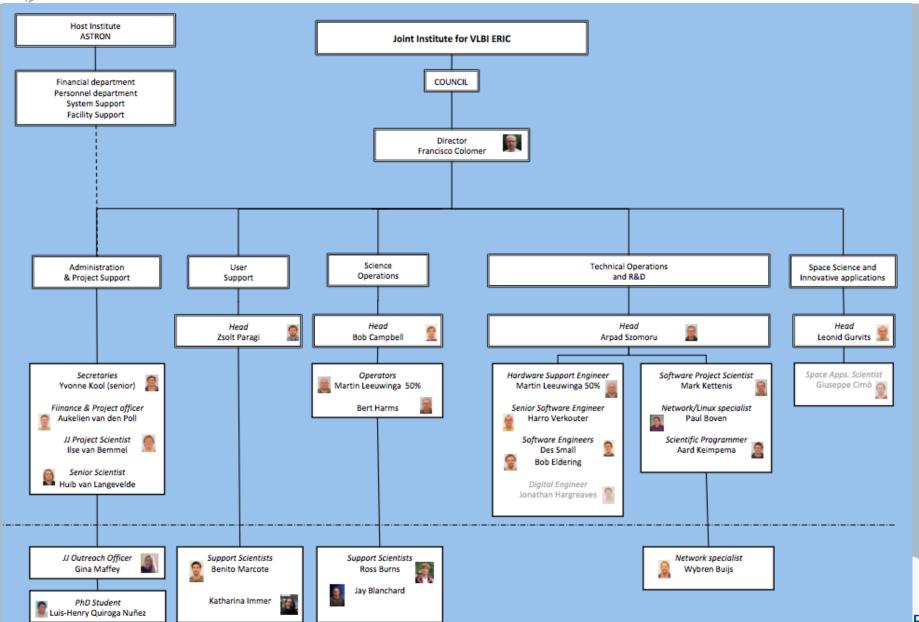
www.jive.eu
@jivevlbi
@JIVERIC

EXTRA SLIDES





Who is who at JIVE





JIVE and EVN

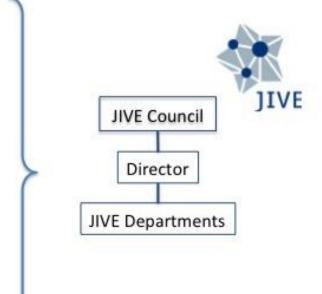
EUROPEAN VLBI NETWORK (EVN)

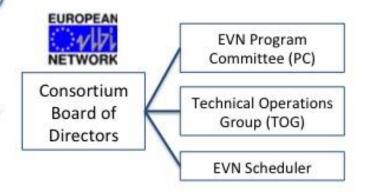
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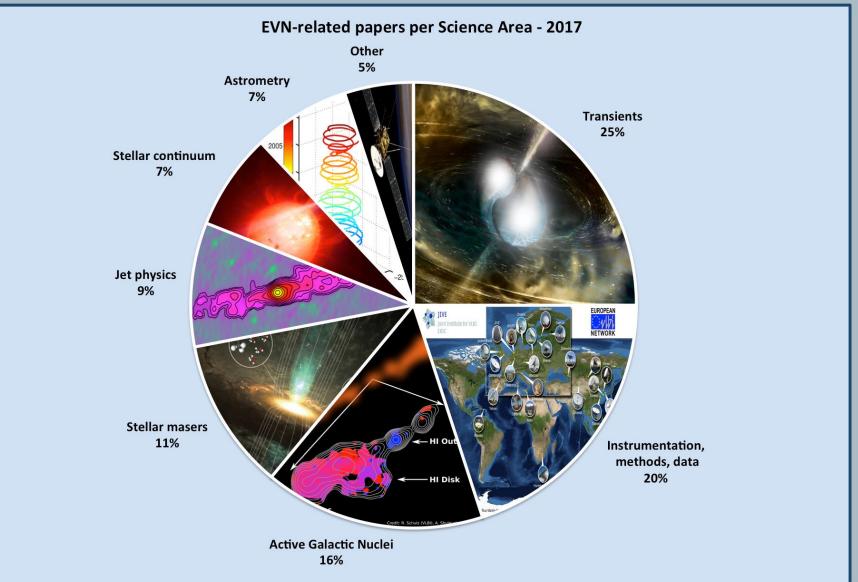








EVN scientific output (2017)





EC Projects

JUMPING JIVE:

- "Joining up Users for Maximizing the Profile, the Innovation and Necessary Globalization of JIVE".
- Coordinated by JIVE
- 3 MEuro in 2017-2020.

RadioNet (EVN TNA, RINGS):

- A new 10 MEuro, 4-year program started in 2017
- TNA provides access to EVN and JIVE.
- RINGS for analysis of wideband data.

ASTERICS:

JIVE coordinates CLEOPATRA and participates in OBELIX











